Spirit Documentation

Release

Gideon Mueller and contributors

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CHAPTER 1

SPIRIT

SPIN SIMULATION FRAMEWORK

Core Library:

Service	System	Compiler	Status
Travis- CI	Ubuntu 14.04 macOS	GCC 6 Clang	master: limage0 ldeve lop: limage1 l
AppVeyo r	Windows	MSVC14 MSVC14.1	master: limage2 ldeve lop: limage3 l

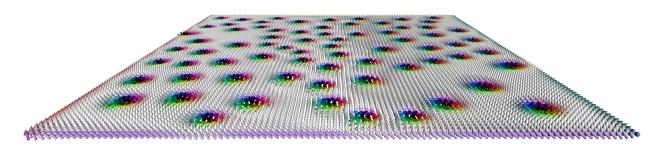
'Python package https://pypi.org/project/spirit/>'__:

Branch	Python Package Coverage	Core Library Coverage
master:		
develop:		

The code is released under MIT License. If you intend to *present and/or publish* scientific results or visualisations for which you used Spirit, please read the REFERENCE.md

This is an open project and contributions and collaborations are always welcome!! See *Contributing* on how to contribute or write an email to g.mueller@fz-juelich.de For contributions and affiliations, see CONTRIBUTORS.md

Please note that a version of the Spirit Web interface is hosted by the Research Centre Jülich at http://juspin.de



1.1 Contents

- 1. Introduction
- 2. Getting started with the Desktop User Interface
- 3. Getting started with the Python Package
- 4. Contributing

1.2 Introduction

1.2.1 A modern framework for magnetism science on clusters, desktops & laptops and even your Phone

Spirit is a **platform-independent** framework for spin dynamics, written in C++11. It combines the traditional cluster work, using using the command-line, with modern visualisation capabilites in order to maximize scientists' productivity.

"It is unworthy of excellent men to lose hours like slaves in the labour of calculation which could safely be relegated to anyone else if machines were used." - Gottfried Wilhelm Leibniz

Our goal is to build such machines. The core library of the *Spirit* framework provides an **easy to use API**, which can be used from almost any programming language, and includes ready-to-use python bindings. A **powerful desktop user interface** is available, providing real-time visualisation and control of parameters.

Physics Features

- Atomistic Spin Lattice Heisenberg Model including also DMI and dipole-dipole
- Spin Dynamics simulations obeying the Landau-Lifschitz-Gilbert equation
- Direct **Energy minimisation** with different solvers
- Minimum Energy Path calculations for transitions between different spin configurations, using the GNEB method

Highlights of the Framework

- Cross-platform: everything can be built and run on Linux, OSX and Windows
- Standalone core library with C API which can be used from almost any programming language
- Python package making complex simulation workflows easy
- Desktop UI with powerful, live 3D visualisations and direct control of most system parameters
- Modular backends including **parallelisation on GPU** (CUDA) and **CPU** (OpenMP)

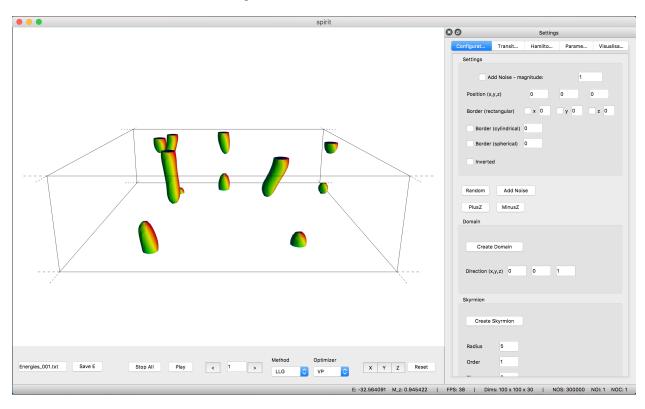
Documentation

More details may be found at spirit-docs.readthedocs.io or in the Reference section including * Framework build instructions * Core build instructions * Core API Reference * Python API Reference * Input File Reference

There is also a Wiki, hosted by the Research Centre Jülich.

1.3 Getting started with the Desktop Interface

See BUILD.md on how to install the desktop user interface.



The user interface provides a powerful OpenGL visualisation window using the VFRendering library. It provides functionality to - Control Calculations - Locally insert Configurations (homogeneous, skyrmions, spin spiral, ...) - Generate homogeneous Transition Paths - Change parameters of the Hamiltonian - Change parameters of the Method and Solver - Configure the Visualization (arrows, isosurfaces, lighting, ...)

See the UI-QT Reference for the key bindings of the various features.

Unfortunately, distribution of binaries for the Desktop UI is not possible due to the restrictive license on QT-Charts.

1.4 Getting started with the Python Package

To install the Spirit python package, either build and install from source or simply use

```
pip install spirit
```

With this package you have access to powerful Python APIs to run and control dynamics simulations or optimizations. This is especially useful for work on clusters, where you can now script your workflow, never having to re-compile when testing, debugging or adding features.

The most simple example of a **spin dynamics simulation** would be

```
from spirit import state, simulation
with state.State("input/input.cfg") as p_state:
    simulation.PlayPause(p_state, "LLG", "SIB")
```

Where "SIB" denotes the semi-implicit method B and the starting configuration will be random.

To add some meaningful content, we can change the **initial configuration** by inserting a Skyrmion into a homogeneous background:

```
def skyrmion_on_homogeneous(p_state):
    from spirit import configuration
    configuration.PlusZ(p_state)
    configuration.Skyrmion(p_state, 5.0, phase=-90.0)
```

If we want to calculate a **minimum energy path** for a transition, we need to generate a sensible initial guess for the path and use the **GNEB method**. Let us consider the collapse of a skyrmion to the homogeneous state:

```
from spirit import state, chain, configuration, transition, simulation
### Copy the system a few times
chain.Image_to_Clipboard(p_state)
for number in range (1,7):
 chain.Insert_Image_After(p_state)
noi = chain.Get_NOI(p_state)
### First image is homogeneous with a Skyrmion in the center
configuration.PlusZ(p_state, idx_image=0)
configuration.Skyrmion(p_state, 5.0, phase=-90.0, idx_image=0)
simulation.PlayPause(p_state, "LLG", "VP", idx_image=0)
### Last image is homogeneous
configuration.PlusZ(p_state, idx_image=noi-1)
simulation.PlayPause(p_state, "LLG", "VP", idx_image=noi-1)
### Create transition of images between first and last
transition. Homogeneous (p_state, 0, noi-1)
### GNEB calculation
simulation.PlayPause(p_state, "GNEB", "VP")
```

where "VP" denotes a direct minimization with the velocity projection algorithm.

You may also use Spirit order to extract quantitative data, such as the energy.

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```
def evaluate(p_state):
    from spirit import system, quantities
    M = quantities.Get_Magnetization(p_state)
    E = system.Get_Energy(p_state)
    return M, E
```

Obviously you may easily create significantly more complex workflows and use Python to e.g. pre- or post-process data or to distribute your work on a cluster and much more!

1.5 Contributing

Contributions are always welcome!

- 1. Fork this repository
- 2. Check out the develop branch: git checkout develop
- 3. Create your feature branch: git checkout -b feature-something
- 4. Commit your changes: git commit -am 'Add some feature'
- 5. Push to the branch: git push origin feature-something
- 6. Submit a pull request

Please keep your pull requests *feature-specific* and limit yourself to one feature per feature branch. Remember to pull updates from this repository before opening a new feature branch.

If you are unsure where to add you feature into the code, please do not hesitate to contact us.

There is no strict coding guideline, but please try to match your code style to the code you edited or to the style in the respective module.

We aim to adhere to the "git flow" branching model: http://nvie.com/posts/a-successful-git-branching-model/

Release versions (master branch) are tagged major.minor.patch, starting at 1.0.0

Download the latest stable version from https://github.com/spirit-code/spirit/releases

The develop branch contains the latest updates, but is generally less consistently tested than the releases.

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Spirit Desktop UI

The cross-platform QT desktop user interface provides a productive tool for Spin simulations, providing powerful real-time visualisations and access to simulation parameters, as well as other very useful features.

See the framework build instructions for information on how to build the user interface on your machine.

2.1 Physics features

Insert Configurations: - White noise - (Anti-) Skyrmions - Domains - Spin Spirals

You may manipulate the Hamiltonian as well as simulation parameters and your output file configuration:

You may start and stop simulation and directly interact with a running simulation. - LLG Simulation: Dynamics and Minimization - GNEB: create transitions and calculate minimum energy paths

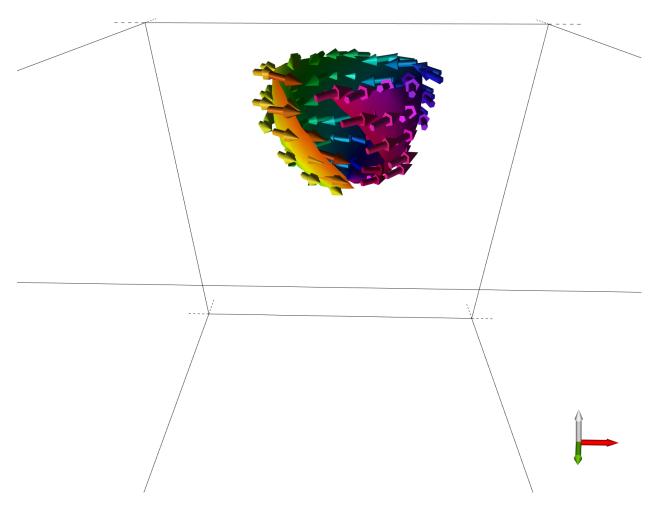
By copying and inserting spin systems and manipulating them you may create arbitrary transitions between different spin states to use them in GNEB calculations. Furthermore you can choose different images to be climbing or falling images during your calculation.

2.2 Real-time visualisation

This feature is most powerful for 3D systems but shows great use for the analysis of dynamical processes and understanding what is happening in your system during a simulation instead of post-processing your data.

- Arrows, Surface (2D/3D), Isosurface
- · Spins or Eff. Field
- · Every n'th arrow
- · Spin Sphere
- Directional & Position filters
- Colormaps

You can also create quite complicate visualisations by combining these different features in order to visualise complex states in 3D systems:



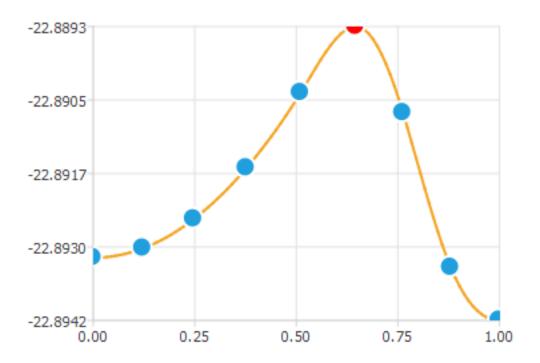
Note that a data plot is available to visualise your chain of spin systems. It can also show interpolated energies if you run a GNEB calculation.

2.3 Additional features

- Drag mode: drag, copy, insert, change radius
- Screenshot
- Read configuration or chain
- Save configuration or chain

2.4 Key bindings

Note that some of the keybindings may only work correctly on US keyboard layout.



2.4.1 UI Controls

Effect	Keystroke
Show this	F1
Toggle Settings	F2
Toggle Plots	F3
Toggle Debug	F4
Toggle "Dragging" mode	F5
Toggle large visualization	F10 / Ctrl+F
Toggle full-screen window	F11 / Ctrl+Shift+F
Screenshot of Visualization region	F12 / Home
Toggle OpenGL Visualization	Ctrl+Shift+V
Try to return focus to main UI	Esc

2.4.2 Camera Controls

Effect	Keystroke
Rotate the camera around	Left mouse / W A S D (Shift to go slow)
Move the camera around	Left mouse / T F G H (Shift to go slow)
Zoom in on focus point	Scroll mouse (Shift to go slow)
Set the camera in X, Y or Z direction	X Y Z (shift to invert)

2.4. Key bindings 9

2.4.3 Control Simulations

Effect	Keystroke
Play/Pause	Space
Cycle Method	Ctrl+M
Cycle Solver	Ctrl+S

2.4.4 Manipulate the current images

Effect	Keystroke
Random configuration	Ctrl+R
Add tempered noise	Ctrl+N
Insert last used configuration	Enter

2.4.5 Visualisation

Effect	Keystroke
Use more/less data points of the vector field	+/-
Regular Visualisation Mode	1
Isosurface Visualisation Mode	2
Slab (X,Y,Z) Visualisation Mode	3 4 5
Cycle Visualisation Mode	1
Move Slab	, / . (Shift to go faster)

2.4.6 Manipulate the chain of images

Effect	Keystroke
Switch between images and chains	\leftarrow \uparrow \rightarrow \downarrow
Cut image	Ctrl+X
Copy image	Ctrl+C
Paste image at current index	Ctrl+V
Insert left/right of current index	$Ctrl+\leftarrow / \rightarrow$
Delete image	Del

Home

Building Spirit's Framework Components

The **Spirit** framework is designed to run across different platforms and so the build process is set up with CMake, which will generate the appropriate build scripts for each platform.

Please be aware that our CMake scripts are written for our use cases and you may need to adapt some paths and options in the Root CMakeLists.txt.

3.1 Contents

- 1. General Build Process
- 2. Core Library
- 3. Desktop User Interface

3.2 General Build Process

The following assumes you are in the Spirit root directory.

3.2.1 Options

There are some important **Options** you may need to consider. You can find them under ### Build Flags ### in the **Root 'CMakeLists.txt <../CMakeLists.txt>'__**. Otherwise, the developers' defaults will be used.

Some Paths you can set under ### User Paths ### (just uncomment the corresponding line) are:

CMake Variable	Use
USER_COMPILER_CUSER_COMPILER_CXX	Name of the compiler you wish to use
USER_PATH_COMPILER	Directory your compiler is located in
USER_PATHS_IFF	use the default IFF (FZJ) cluster paths

3.2.2 Clean

Clear the build directory using

```
./clean.sh
or
rm -rf build && mkdir build
```

Further helper scripts for clean-up are clean_log.sh, clean_output.sh,

3.2.3 Generate Build Files

./cmake.sh lets cmake generate makefiles for your system inside a 'build' folder. Simply call

```
./cmake.sh
or
cd build && cmake .. && cd ..
```

Passing -debug to the script will cause it to create a debug configuration, meaning that you will be able to properly debug the entire application.

On **Windows** (no MSys) you can simply use the git bash to do this or use the CMake GUI. When using MSys etc., CMake will create corresponding MSys makefiles.

3.2.4 Building the Projects

To execute the build and linking of the executable, simply call

```
./make.sh -jN or cd build && make -jN && cd ..
```

where - jN is optional, with N the number of parallel build processes you want to create.

On **Windows** (no MSys), CMake will by default have generated a Visual Studio Solution. Open the generated Solution in the Visual Studio IDE and build it there.

3.2.5 Running the Unit Tests

We use CMakes CTest for unit testing. You can run

```
ctest.sh
or
cd build && ctest --output-on-failure && cd ..
```

or execute any of the test executables manually. To execute the tests from the Visual Studio IDE, simply rebuild the RUN_TESTS project.

3.2.6 Installing Components

This is not yet supported! however, you can already run

```
./install.sh
or
cd build && make install && cd ..
```

Which on OSX should build a .app bundle.

3.3 Core Library

For detailed build instructions concerning the standalone core library or how to include it in your own project, see core/docs/BUILD.md. * Shared and static library * Python bindings * Julia bindings * Transpiling to JavaScript * Unit Tests

The Root 'CMakeLists.txt <../CMakeLists.txt>'__ has a few options you can set:

CMake Options	Use
SPIRIT_USE_CUDA Use CUDA to spee d up nume rically intensive part s of the core	
SPIRIT_USE_OPENMP	Use Open MP to spee d up nume rica lly inte nsiv e part s of the core
SPIRIT_SCALAR_TYPE	Shou ld be e.g. "do uble " or "fl oat". Sets the C++ type for scal ar vari able s,
	arra ys etc.
SPIRIT_BUILD_TEST	Buil d unit test s for the core libr ary
SPIRIT_BUILD_FOR_CXX	Buil d the stat ic libr ary for C++ appl icat ions
SPIRIT_BUILD_FOR_JULI	Buil d the shar ed libr ary for Juli a
A	
SPIRIT_BUILD_FOR_PYTH	Buil d the shar ed libr ary for Pyth on
ON	
SPIRIT_BUILD_FOR_JS	Buil d the Java Scri pt libr ary (use s a diff eren t tool chai n!)

3.4 Desktop User Interface

Dependencies	Versions
OpenGL Drivers	>= 3.3
CMake	>= 3.1
QT	>= 5.7 including QT-Charts

Note that in order to build with QT as a dependency on Windows, you may need to add path/to/qt/qtbase/bin to your PATH variable.

Necessary OpenGL drivers should be available through the regular drivers for any remotely modern graphics card.

CMake Options	Use
SPIRIT_BUILD_FOR_C XX	Buil d the C++ inte rfac es (con sole or QT) inst ead of othe rs
UI_CXX_USE_QT	Buil d qt user inte rfac e inst ead of cons ole vers ion
USER_PATH_QT	The path to your CMak e inst alla tion
BUNDLE_APP	On OSX, crea te .app bund le (not yet full y func tion al)

3.3. Core Library

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CHAPTER 4

SPIRIT INPUT FILES

The following sections will list and explain the input file keywords.

- 1. General Settings and Log
- 2. Geometry
- 3. Hamiltonian
- 4. Method Output
- 5. Method Parameters
- 6. Pinning
- 7. Disorder and Defects

4.1 General Settings and Log

```
### Save input parameters on creation of State
log_input_save_initial 0
### Save input parameters on deletion of State
log_input_save_final 0

### Print log messages to the console
log_to_console 1
### Print messages up to (including) log_console_level
log_console_level 5

### Save the log as a file
log_to_file 1
### Save messages up to (including) log_file_level
log_file_level 5
```

Except for SEVERE and ERROR, only log messages up to log_console_level will be printed and only messages up to log_file_level will be saved. If log_to_file, however is set to zero, no file is written at all.

Log Levels	Integer	Description
ALL	0	Everything
SEVERE	1	Only severe errors
ERROR	2	Also non-fatal errors
WARNING	3	Also warnings
PARAMETER	4	Also input parameters
INFO	5	Also info-messages
DEBUG	6	Also deeper debug-info

4.2 Geometry

The Geometry of a spin system is specified in form of a bravais lattice and a basis cell of atoms. The number of basis cells along each principal direction of the basis can be specified. *Note:* the default basis is a single atom at (0,0,0).

3D simple cubic example:

```
### The bravais lattice type
bravais_lattice sc

### Number of basis cells along principal
### directions (a b c)
n_basis_cells 100 100 10
```

2D honeycomb example:

```
### The bravais lattice type
bravais_lattice hex2d

### n          No of spins in the basis cell
### 1.x 1.y 1.z position of spins within basis
### 2.x 2.y 2.z cell in terms of bravais vectors
basis
2
0 0 0 0
0.86602540378443864676 0.5 0

### Number of basis cells along principal
### directions (a b c)
n_basis_cells 100 100 1
```

The bravais lattice can be one of the following:

Bravais Lattice Type	Keyword	Comment
Simple cubic	sc	
Body-centered cubic	bcc	
Face-centered cubic	fcc	
Hexagonal (2D)	hex2d	60deg angle
Hexagonal (2D)	hex2d60	60deg angle
Hexagonal (2D)	hex2d120	120deg angle
Hexagonal closely packed	hcp	120deg, not yet implemented
Hexagonal densely packed	hdp	60deg, not yet implemented
Rhombohedral	rho	not yet implemented
Simple-tetragonal	stet	not yet implemented
Simple-orthorhombic	so	not yet implemented
Simple-monoclinic	sm	not yet implemented
Simple triclinic	stri	not yet implemented

Alternatively it can be input manually, either through vectors or as the bravais matrix:

```
### bravais_vectors or bravais_matrix
### a.x a.y a.z a.x b.x c.x
### b.x b.y b.z a.y b.y c.y
### c.x c.y c.z a.z b.z c.z
bravais_vectors
1.0 0.0 0.0
0.0 1.0 0.0
0.0 0.0 1.0
```

A lattice constant can be used for scaling:

```
### Scaling constant
lattice_constant 1.0
```

4.3 Hamiltonian

Note that you select the Hamiltonian you use with the hamiltonian keyword.

Isotropic Heisenberg Hamiltonian:

Interactions are handled in terms of neighbours. You may specify shell-wise interaction parameters:

```
### Hamiltonian Type (heisenberg_neighbours, heisenberg_pairs, gaussian)
hamiltonian
                        heisenberg_neighbours
### boundary_conditions (in a b c) = 0(open), 1(periodical)
boundary_conditions
                       1 1 0
### external magnetic field vector[T]
external_field_magnitude 25.0
external_field_normal 0.0 0.0 1.0
### µSpin
                        2.0
mu_s
### Uniaxial anisotropy constant [meV]
anisotropy_magnitude 0.0
anisotropy_normal
                        0.0 0.0 1.0
```

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If you have a nontrivial basis cell, note that you should specify mu_s for all atoms in your basis cell.

Anisotropy: By specifying a number of anisotropy axes via n_anisotropy, one or more anisotropy axes can be set for the atoms in the basis cell. Specify columns via headers: an index i and an axis Kx Ky Kz or Ka Kb Kc, as well as optionally a magnitude K.

Pair-wise Heisenberg Hamiltonian:

Interactions are specified pair-wise. Single-threaded applications can thus calculate interactions twice as fast as for the neighbour-wise case. You may specify shell-wise interaction parameters.

```
### Hamiltonian Type (heisenberg_neighbours, heisenberg_pairs, gaussian)
                       heisenberg_pairs
### Boundary_conditions (in a b c) = 0(open), 1(periodical)
boundary_conditions
                 1 1 0
### External magnetic field vector[T]
external_field_magnitude 25.0
external_field_normal
                       0.0 0.0 1.0
### µSpin
                        2.0
mu_s
### Uniaxial anisotropy constant [meV]
anisotropy_magnitude 0.0
anisotropy_normal
                      0.0 0.0 1.0
### Dipole-Dipole radius
                       0.0
dd_radius
### Pairs
n_interaction_pairs 3
i j da db dc Jij Dij Dijx Dijy Dijz
   1 0 0
              10.0 6.0 1.0 0.0 0.0
0 0
     0 1 0 10.0 6.0 0.0 1.0 0.0
    0 0 1 10.0 6.0
0 0
                        0.0 0.0 1.0
### Quadruplets
n_interaction_quadruplets 1
   j da_j db_j dc_j k da_k db_k dc_k l da_l db_l dc_l
                                                               Q
    0 1
           0 0 0 0 1 0 0 0
                                                   0
                                                        1
                                                               3.0
```

If you have a nontrivial basis cell, note that you should specify mu_s for all atoms in your basis cell.

Anisotropy: By specifying a number of anisotropy axes via n_anisotropy, one or more anisotropy axes can be set for the atoms in the basis cell. Specify columns via headers: an index i and an axis Kx Ky Kz or Ka Kb Kc, as well as optionally a magnitude K.

Pairs: Leaving out either exchange or DMI in the pairs is allowed and columns can be placed in arbitrary order. Note that instead of specifying the DM-vector as Dijx Dijy Dijz, you may specify it as Dija Dijb Dijc if you prefer. You may also specify the magnitude separately as a column Dij, but note that if you do, the vector (e.g. Dijx Dijy Dijz) will be normalized.

Quadruplets: Columns for these may also be placed in arbitrary order.

Separate files: The anisotropy, pairs and quadruplets can be placed into separate files, you can use anisotropy_from_file, pairs_from_file and quadruplets_from_file.

If the headers for anisotropies, pairs or quadruplets are at the top of the respective file, it is not necessary to specify n_anisotropy, n_interaction_pairs or n_interaction_quadruplets respectively.

```
### Pairs
interaction_pairs_file input/pairs.txt

### Quadruplets
interaction_quadruplets_file input/quadruplets.txt
```

Gaussian Hamiltonian:

This is a testing Hamiltonian consisting of the superposition of gaussian potentials. It does not contain interactions.

```
hamiltonian gaussian

### Number of Gaussians
n_gaussians 2

### Gaussians
### a is the amplitude, s is the width, c the center
### the directions c you enter will be normalized
### a1 s1 c1.x c1.y c1.z
### ...
gaussians
1 0.2 -1 0 0
0.5 0.4 0 0 -1
```

4.4 Method Output

For 11g and equivalently mc and gneb, you can specify which output you want your simulations to create. They share a few common output types, for example:

```
      llg_output_any
      1
      # Write any output at all

      llg_output_initial
      1
      # Save before the first iteration

      llg_output_final
      1
      # Save after the last iteration
```

Note in the following that step means after each N iterations and denotes a separate file for each step, whereas archive denotes that results are appended to an archive file at each step.

LLG:

```
      llg_output_energy_step
      0
      # Save system energy at each step

      llg_output_energy_archive
      1
      # Archive system energy at each step
```

```
      llg_output_energy_spin_resolved
      0
      # Also save energies for each spin

      llg_output_energy_divide_by_nspins
      1
      # Normalize energies with number of spins

      llg_output_configuration_step
      1
      # Save spin configuration at each step

      llg_output_configuration_archive
      0
      # Archive spin configuration at each step
```

MC:

```
mc_output_energy_step 0
mc_output_energy_archive 1
mc_output_energy_spin_resolved 0
mc_output_energy_divide_by_nspins 1
mc_output_configuration_step 1
mc_output_configuration_archive 0
```

GNEB:

4.5 Method Parameters

Again, the different Methods share a few common parameters. On the example of the LLG Method:

LLG:

```
### Spin transfer torque parameter proportional to injected current density
llg_stt_magnitude 0.0
### Spin current polarisation normal vector
llg_stt_polarisation_normal 1.0 0.0 0.0
```

MC:

```
### Seed for Random Number Generator
mc_seed 20006

### Temperature [K]
mc_temperature 0

### Acceptance ratio
mc_acceptance_ratio 0.5
```

GNEB:

```
### Constant for the spring force
gneb_spring_constant 1.0

### Number of energy interpolations between images
gneb_n_energy_interpolations 10
```

4.6 Pinning

Note that for this feature you need to build with SPIRIT_ENABLE_PINNING set to ON in cmake.

For each lattice direction a b and c, you have two choices for pinning. For example to pin n cells in the a direction, you can set both pin_na_left and pin_na_right to different values or set pin_na to set both to the same value. To set the direction of the pinned cells, you need to give the pinning_cell keyword and one vector for each basis atom.

You can for example do the following to create a U-shaped pinning in x-direction:

```
# Pin left side of the sample (2 rows)
pin_na_left 2
# Pin top and bottom sides (2 rows each)
pin_nb 2
# Pin the atoms to x-direction
pinning_cell
1 0 0
```

To specify individual pinned sites (overriding the above pinning settings), insert a list into your input. For example:

```
### Specify the number of pinned sites and then the directions
### ispin S_x S_y S_z
n_pinned 3
0 1.0 0.0 0.0
1 0.0 1.0 0.0
2 0.0 0.0 1.0
```

You may also place it into a separate file with the keyword pinned_from_file, e.g.

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```
### Read pinned sites from a separate file
pinned_from_file input/pinned.txt
```

The file should either contain only the pinned sites or you need to specify n_pinned inside the file.

4.7 Disorder and Defects

Note that for this feature you need to build with SPIRIT_ENABLE_DEFECTS set to ON in cmake.

Disorder is not yet implemented.

To specify defects, be it vacancies or impurities, you may fix atom types for sites of the whole lattice by inserting a list into your input. For example:

```
### Atom types: type index 0..n or or vacancy (type < 0)
### Specify the number of defects and then the defects
### ispin itype
n_defects 3
0 -1
1 -1
2 -1</pre>
```

You may also place it into a separate file with the keyword defects_from_file, e.g.

```
### Read defects from a separate file defects_from_file input/defects.txt
```

The file should either contain only the defects or you need to specify n_defects inside the file.

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BUILD THE SPIRIT LIBRARY

5.1 C/C++ Library

Dependencies	Versions
gcc	>= 5.1 (C++11 stdlib) or any modern compiler
CMake	>= 3.1

This is pretty much a standalone library and should be easy to implement into existing projects in which CMake is already used. It can be built as shared or static and API headers are located in path/to/Spirit/core/include/Spirit

CMake Options	Use
SPIRIT_USE_CUDA	Use CUDA to spee d up nume rica lly inte nsiv e part s of the core
SPIRIT_USE_OPENMP	Use Open MP to spee d up nume rica lly inte nsiv e part s of the core
SPIRIT_SCALAR_TYPE	Shou ld be e.g. "do uble " or "fl oat". Sets the C++ type for scal ar vari able s,
	arra ys etc.
SPIRIT_BUILD_TEST	Buil d unit test s for the core libr ary
SPIRIT_BUILD_FOR_CXX	Buil d the stat ic libr ary for C++ appl icat ions
SPIRIT_BUILD_FOR_JULI	Buil d the shar ed libr ary for Juli a
A	
SPIRIT_BUILD_FOR_PYTH	Buil d the shar ed libr ary for Pyth on
ON	
SPIRIT_BUILD_FOR_JS	Buil d the Java Scri pt libr ary (use s a diff eren t tool chai n!)

Note that the CMake Options for the core library are also set in the root CMakeLists.txt.

5.1.1 Integrating into other CMake Projects

This is currently untested, but you should be able to use *Spirit* either via ExternalProject or AddSubdirectory. You should thus be able to simply copy the core directory as Spirit into the thirdparty

folder of your Project and use it as is done commonly.

Note that setting <code>qhull_LIBS</code> and <code>qhull_INCLUDE_DIRS</code> if <code>qhull</code> is present on your disk removes the need for Spirit to clone and build it separately.

A set of CMake variables is pushed to the parent scope by the core CMakeLists. These include: - SPIRIT_LIBRARIES - SPIRIT_LIBRARIES_STATIC - SPIRIT_INCLUDE_DIRS

5.2 Unit Tests

Dependencies	Versions
gcc	>= 4.8.1 (C++11 stdlib) or any modern compiler
CMake	>= 2.8.12
Python	>= 2.7 or >= 3.5 (for the python tests)

The core library includes a set of unit tests in order to make sure certain conditions are fulfilled by the librarys functions. We use CMakes CTest for unit testing. You can run

```
cd build && ctest --output-on-failure && cd ..
```

or execute any of the test executables manually. To execute the tests from the Visual Studio IDE, simply rebuild the RUN_TESTS project.

CMake Options	Use
CORE_BUILD_TEST	Build unit tests for the core library
BUILD_UI_PYTHON	Python tests are only built if this is ON

5.3 Python Package

Dependencies	Versions
Python	>= 2.7 or >= 3.5

CMake Options	Use
BUILD_UI_PYTHON	Buil d the shar ed libr ary for pyth on
CORE_BUILD_TEST	Will buil d pyth on unit test s if the shar ed lib for pyth on is buil t

You may use the python package simply by adding a path/to/Spirit/core/python to your PYTHONPATH.

To build a proper python package including binary wheel, you need to build the shared library for Python and then execute

```
cd path/to/Spirit/core/python
  python setup.py sdist bdist_wheel
```

The resulting package can then be installed with e.g.

pip install -e /path/to/package/dist/spirit-<tags>.whl

5.4 Julia Package

These bindings are currently completely experimental. However, it seems it would be easy to create Julia bindings analogous to the *Python Bindings*.

Dependencies	Versions
None	•

CMake Options	Use
BUILD_UI_JULIA	Build the shared library for julia

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CHAPTER 6

SPIRIT API

This will list the available API functions of the Spirit library.

The API is exposed as a C interface and may thus also be used from other languages, such as Python, Julia or even JavaScript (see *ui-web*). The API revolves around a simulation State which contains all the necessary data and keeps track of running Solvers.

The API exposes functions for: * Control of simulations * Manipulation of parameters * Extracting information * Generating spin configurations and transitions * Logging messages * Reading spin configurations * Saving datafiles

6.1 State Managment

To create a new state with one chain containing a single image, initialized by an input file, and run the most simple example of a **spin dynamics simulation**:

A new state can be created with State_Setup(), where you can pass a config file specifying your initial system parameters. If you do not pass a config file, the implemented defaults are used. Note that you currently cannot change the geometry of the systems in your state once they are initialized.

The State struct is passed around in an application to make the simulation's state available.

State manipulation function	Retur	Effe ct
	n	
State_Setup(const char * config_file)	"Sta te	Crea te new stat e by pass ing a conf ig
	*''	file
State_Update(State *)	voi d	Upda te the stat e to hold curr ent valu
		es
State_Delete(State *)	voi d	Dele te a stat e
State_To_Config(State *, const	voi d	Writ e a conf ig file whic h will result in
<pre>char * config_file, const ch ar *</pre>		the same stat e if used in "St ate_ Setu
original_config_file)		p()' '
State_DateTime(State *)	Saan at	Get date time tag of the crea tion of the
	"con st	stat e
	cii ar *	
	-	

6.2 System

System Information	Re-	Effect
	turn	
System_Get_Index(State *)	int	Returns System's Index
System_Get_NOS(State *, int idx_image , int	int	Return System's number of
idx_chain)		spins

System Data	Re-	Effe ct
	turn	
System_Get_Spin_Directions(State *, int	scal	Get Syst em's spin dire ctio
<pre>idx_image, int id x_chain)</pre>	ar *	n
System_Get_Spin_Effective_Field(State *, int	scal	Get Syst em's spin effe ctiv
<pre>idx_image, i nt idx_chain)</pre>	ar *	e fiel d
System_Get_Rx(State *, int idx_image, int		Get a Syst em's reac tion
idx_chain)	t	coor dina te in it's chai n
"System_Get_Energy(State *, int idx_image, int idx_chain) "		Get Syst em's ener gy
<pre>System_Get_Energy_Array(State * energies, float * energie s, int idx_image, int idx_chain)</pre>	"void	Ener gy Arra y (Sho uld NOT be used)

System Output	Effect
System_Print_Energy_Array(State *, int idx_image,	Print on the consol e State's en-
int idx_chai n)	ergy array

System Update	Effect
System_Update_Data(State *, int idx_image, int	Update State's data. Used mainly for
idx_chain)	plotti ng

6.3 Simulation

 $With \ {\tt Simulation_*} \ functions \ one \ can \ control \ and \ get \ information \ from \ the \ State's \ Solvers \ and \ their \ Methods.$

Simulation Basics	Effect
Simulation_SingleShot(State *, const char *	Execut es a single Optimi zation iterat
<pre>c_method_type, con st char * c_solver_type,</pre>	ion with a given method (CAUTI ON!
<pre>int n_iterations, int n_iterations_log, int</pre>	does not check for alread y runnin g
<pre>idx_image, int idx_chain)</pre>	simula tions)
Simulation_PlayPause(State *, const char *	Play/P ause functi onalit y
<pre>c_method_type, cons t char * c_solver_type,</pre>	
<pre>int n_iterations, int n_iterations_log, i nt</pre>	
<pre>idx_image, int idx_chain)</pre>	
Simulation_Stop_All(State *)	Stop all State's simula tions

Simulation Data	Return	Effe ct
Simulation_Get_MaxTorqueComponent(State *,	float	Get Simu lati on's maxi mum
int idx_i mage, int idx_chain)		torq ue comp onen t
Simulation_Get_IterationsPerSecond(State *,	float	Get Simu lati on's iter atio ns
int idx_ image, int idx_chain)		per seco nd
Simulation_Get_Solver_Name(State *, int	const	Get Solv er's name
idx_image, i nt idx_chain)	cha r *	
Simulation_Get_Method_Name(State *, int	const	Get Meth od's name
idx_image, i nt idx_chain)	cha r *	

Simulation Running Checking	Return
Simulation_Running_Any_Anywhere(State *)	
Simulation_Running_LLG_Anywhere(State *)	
Simulation_Running_GNEB_Anywhere(State *)	
Simulation_Running_LLG_Chain(State *state, int idx_chain)	
Simulation_Running_Any(State *, int idx_image, int idx_chain)	
Simulation_Running_LLG(State *, int idx_image, int idx_chain)	
Simulation_Running_GNEB(State *, int idx_chain)	
Simulation_Running_MMF(State *)	

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6.4 Geometry

Geometry Data	Return
<pre>Geometry_Get_Spin_Positions(State *,</pre>	scala r *
<pre>int idx_image, int idx_c hain)</pre>	
Geometry_Get_Bounds(State *, float	66 . 16 6
<pre>min[3], float max[3], int idx_image,</pre>	"void"
int idx_chain)	
<pre>Geometry_Get_Center(State *,</pre>	"void"
<pre>float center[3], int idx_image, int</pre>	Void
idx_chain)	
"Geometry_Get_Cell_Bounds(State *, float min[3], f	loat-max[3].
int idx_image, int idx_chain)"	1010 1 37
<pre>Geometry_Get_Basis_Vectors(State *,</pre>	int
float a[3], float b[3], f loat c[3],	
<pre>int idx_image, int idx_chain)</pre>	
<pre>Geometry_Get_N_Basis_Atoms(State *,</pre>	66 * 16 6
<pre>int idx_image, int idx_ch ain)</pre>	"void"
<pre>Geometry_Get_N_Cells(State *,</pre>	"void"
<pre>int n_cells[3], int idx_image, int</pre>	Void
idx_chain)	
Geometry_Get_Translation_Vectors("void"
State *, float ta[3], float tb[3],	Void
<pre>float tc[3], int idx_image, int</pre>	
idx_chain)	
Geometry_Get_Dimensionality(State *,	int
<pre>int idx_image, int idx_c hain)</pre>	
Geometry_Get_Triangulation(State	int
*, const int **indices_ptr, int	
n_cell_step, int idx_image, int	
idx_chain)	

6.5 Transitions

Transitions	Return	Effect
"Transition_Homogeneous(State * int idx chain=-1)"	, int idx_1, int idx_2,	•
_ ,		
Transition_Add_Noise_Temp	e va tidre(•
State *, float tempe		
rature, int idx_1, int		
idx_2, int idx_chain)		

6.6 Quantities

Quantities	Return	Effect
Quantity_Get_Magnetizatio	n∜oid	•
State *, float m[3],		
int i dx_image, int		
idx_chain)		
Quantity_Get_Topological_	Charge (•
State *, int idx_imag e,	"float"	
int idx_chain)		

6.7 Parameters

LLG Parameters Set	Retu rn	Effect
Parameters_Set_LLG_Time_S	t wp (id	•
State *, float dt,		
int idx_i mage, int		
idx_chain)		
Parameters_Set_LLG_Dampin	g∜o id	•
State *, float damping,		
int id x_image, int		
idx_chain)		
Parameters_Set_LLG_N_Iter	atrioonisi(•
State *, int		
n_iterations , int		
idx_image, int idx_chain		
)		

GNEB Parameters Set	Retu rn	Effect
Parameters_Set_GNEB_Sprin	g <u>v</u> 6omistant (•
State *, float		
spring _constant, int		
idx_image, int idx_chain		
)		
Parameters_Set_GNEB_Climb	ixg_falling(•
State *, int image_t		
ype, int idx_image, int		
idx_chain)		
Parameters_Set_GNEB_N_Ite	ræbions(•
State *, int n_iteration		
s, int idx_chain)		

6.6. Quantities 31

LLG Parameters Get	Retu rn	Effect
Parameters_Get_LLG_Time_S	t ep (id	•
State *, float * dt,		
int idx _image, int		
idx_chain)		
Parameters_Get_LLG_Dampin	g∜o id	•
State *, float *		
damping, int idx_image,		
int idx_chain)		
Parameters_Get_LLG_N_Iter	atroonid(•
State *, int idx_image,		
i nt idx_chain)		

GNEB Parameters Get	Retu rn	Effect
Parameters_Get_GNEB_Sprin	g <u>v</u> 6omisitant (•
State *, float *		
spri ng_constant, int		
idx_image, int idx_chain		
)		
Parameters_Get_GNEB_Climb	i xg_fa lling(•
State *, int * image		
_type, int idx_image,		
int idx_chain)		
"Parameters_Get_GNEB_N_Iteration State *, int idx_chain) "	in t ns(•
<pre>Parameters_Get_GNEB_N_Ene State *, int idx chain)</pre>	rġņ_Enterpolations(•
beace , inc iax_chain ,		

6.8 Chain

Get Chain's information

Chain Info	Return	Effect
<pre>Chain_Get_Index(State *)</pre>	int	•
Chain_Get_NOI(State *, int idx_chain)	int	•

Move between images (change $active_image$)

Chain moving		Return
Chain_prev_Image(State *, int idx_chain)		bool
Chain_next_Image(State *, int idx_chain)		bool
Chain_Jump_To_Image(State *, int idx_image, int idx_	_chain)	bool

Insertion/deletion and replacement of images are done by

Chain control	Return
Chain_Image_to_Clipboard(State *, int idx_image, int idx_chain)	void
Chain_Replace_Image(State *, int idx_image, int idx_chain)	void
Chain_Insert_Image_Before(State *, int idx_image, int idx_chain)	void
Chain_Insert_Image_After(State *, int idx_image, int idx_chain)	void
Chain_Push_Back(State *, int idx_chain)	void
Chain_Delete_Image(State *, int idx_image, int idx_chain)	bool
Chain_Pop_Back(State *, int idx_chain)	bool

Chain data	Re-
	turn
Chain_Get_Rx(State *, float* Rx, int idx_chain)	"void
<pre>Chain_Get_Rx_Interpolated(State *, float * Rx_interpolated, in t idx_chain)</pre>	"void
Chain_Get_Energy(State *, float* energy, int idx_chain)	"void
<pre>Chain_Get_Energy_Interpolated(State *, float* E_interpolated, int idx_chain)</pre>	"void

Chain data update	Return
Chain_Update_Data(State *, int idx_chain)	void
Chain_Setup_Data(State *, int idx_chain)	void

6.9 Hamiltonian

Set Hamiltonian's parameters

6.9. Hamiltonian 33

Hamiltonian Set	Re-
	tur
	n
<pre>Hamiltonian_Set_Boundary_Conditions(State *, const bool* period ical,</pre>	voi
<pre>int idx_image, int idx_chain)</pre>	d
Hamiltonian_Set_mu_s(State *, float mu_s, int idx_image, int id	voi
x_chain)	d
Hamiltonian_Set_Field(State *, float magnitude, const float* no rmal,	voi
<pre>int idx_image, int idx_chain)</pre>	d
Hamiltonian_Set_Exchange(State *, int n_shells, const float* ji j,	voi
<pre>int idx_image, int idx_chain)</pre>	d
Hamiltonian_Set_DMI(State *, float dij, int idx_image, int idx_ chain	voi
	d
<pre>Hamiltonian_Set_BQE(State *, float dij, int idx_image, int idx_ chain</pre>	voi
	d
Hamiltonian_Set_FSC(State *, float dij, int idx_image, int idx_ chain	voi
	d
Hamiltonian_Set_Anisotropy(State *, float magnitude, const floa t*	voi
normal, int idx_image, int idx_chain)	d
<pre>Hamiltonian_Set_STT(State *, float magnitude, const float * nor mal,</pre>	voi
<pre>int idx_image, int idx_chain)</pre>	d
Hamiltonian_Set_Temperature(State *, float T, int idx_image, in t	voi
idx_chain)	d

Get Hamiltonian's parameters

Hamiltonian Get	Return
Hamiltonian_Get_Name(State *, int	const c har *
<pre>idx_image, int idx_chain)</pre>	
Hamiltonian_Get_Boundary_Conditions(void
State *, bool* periodic al, int	
<pre>idx_image, int idx_chain)</pre>	
<pre>Hamiltonian_Get_mu_s(State *, float*</pre>	void
<pre>mu_s, int idx_image, i nt idx_chain)</pre>	
<pre>Hamiltonian_Get_Field(State *,</pre>	void
float* magnitude, float* nor mal,	
<pre>int idx_image, int idx_chain)</pre>	
Hamiltonian_Get_Exchange(State	void
, int n_shells, float* jij , int	
<pre>idx_image, int idx_chain)</pre>	
<pre>Hamiltonian_Get_Anisotropy(State *,</pre>	void
<pre>float* magnitude, float * normal, int</pre>	
<pre>idx_image, int idx_chain)</pre>	
"Hamiltonian_Get_DMI(State , float dij, int idx_ima idx_chain)"	void ge, int
"Hamiltonian_Get_BQE(State , float dij, int idx_ima idx_chain)"	void ge, int
"Hamiltonian_Get_FSC(State, float dij, int idx_ima idx_chain)"	void ge, int
<pre>Hamiltonian_Get_STT(State *, float*</pre>	void
magnitude, float* norma l, int	
idx_image, int idx_chain)	
Hamiltonian_Get_Temperature(State	void
, float T, int idx_imag e, int	
idx_chain)	

6.10 Constants

Constants	Return	Description
Constants_mu_B()	scalar	The Bohr Magneton [meV / T]
Constants_k_B()	scalar	The Boltzmann constant [meV / K]

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6.11 Log

Log Utilities	Retur n	Effect
Log_Send(State *, int level, int	voi d	Send a Log message
sender, const char * m essage, int		
<pre>idx_image, int idx_chain)</pre>		
Log_Get_Entries(State *)	std ::vec	Get the entries from the Log
	tor <u td="" tilit<=""><td>and write new number of en-</td></u>	and write new number of en-
	y::Lo gEntr y>	tries into given int
Log_Append(State *)	voi d	Append the Log to it's file
Log_Dump(State *)	voi d	Dump the Log into it's file
Log_Get_N_Entries(State *)	"int "	Get the number of Log entries
Log_Get_N_Errors(State *)	"int "	Get the number of errors in
	- IIIt	the Log
Log_Get_N_Warnings(State *)	66:	Get the number of warning s
	"int "	in the Log

Log macro variables for Levels

Log Levels	value
Log_Level_All	0
Log_Level_Severe	1
Log_Level_Error	2
Log_Level_Warning	3
Log_Level_Parameter	4
Log_Level_Info	5
Log_Level_Debug	6

Log macro variables for Senders

Log Senders	value
Log_Sender_All	0
Log_Sender_IO	1
Log_Sender_GNEB	2
Log_Sender_LLG	3
Log_Sender_MC	4
Log_Sender_MMF	5
Log_Sender_API	6
Log_Sender_UI	7

6.12 IO

Macros of File Formats for Vector Fields	values	Description
IO_Fileformat_Regular	0	sx sy sz (separated by whitespace)
IO_Fileformat_Regular_Pos	1	px py pz sx sy sz (separated by whitespace)
IO_Fileformat_CSV	2	sx, sy, sz (separated by commas)
IO_Fileformat_CSV_Pos	3	px, py, pz, sx, sy, (sz separated by commas)
IO_Fileformat_OVF_bin8	4	OOMMF vector field (OVF) v2.0 file format
IO_Fileformat_OVF_text	6	

Read and Write functions

Images IO	Retur n
"IO_N_Images_In_File(State *state, const char *file, int idx_chain)"	int format,
IO_Image_Write(State *state, const	voi d
char *file, int format, cons t	
char *comment, int idx_image, int	
idx_chain)	
IO_Image_Append(State *state,	voi d
const char *file, int format, con	
st char *comment, int idx_image, int	
idx_chain)	
IO_Image_Read(State *state,	voi d
const char *file, int format,	
<pre>int i dx_image_infile, int</pre>	
<pre>idx_image_inchain, int idx_chain)</pre>	

Chains IO	Re-
	tur
	n
<pre>IO_Chain_Write(State *state, const char *file, int format, con st char</pre>	voi
*comment, int idx_chain)	d
<pre>IO_Chain_Append(State *state, const char *file, int format, con st</pre>	voi
<pre>char *comment, int idx_chain)</pre>	d
<pre>IO_Chain_Read(State *state, const char *file, int format, int s</pre>	voi
tarting_image, int ending_image, int insert_idx, int idx_chain)	d

Energies from System and Chain

System Energies	Return
"IO_Image_Write_Energy_per_Spin(State, const chidx_image, int idx_chain)"	<i>ar</i> ંધ્રીલ _{દી} int
<pre>IO_Image_Write_Energy(State *, const char* file, int idx_imag e, int idx_chain)</pre>	"void"

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Chain Energies	Re-
	turn
<pre>IO_Chain_Write_Energies(State *, const char* file, int idx_ch ain)</pre>	"void"
<pre>IO_Chain_Write_Energies_Interpolated(State *, const char* fil e, int idx_chain)</pre>	"void"

Home

CHAPTER 7

SPIRIT Python API

7.1 State

To create a new state with one chain containing a single image, initialized by an input file, and run the most simple example of a **spin dynamics simulation**:

```
from spirit import state
from spirit import simulation

cfgfile = "input/input.cfg"  # Input File
with state.State(cfgfile) as p_state:  # State setup
    simulation.PlayPause(p_state, "LLG", "SIB") # Start a LLG simulation using the_
→SIB solver
```

or call setup and delete manually:

You can pass a config file specifying your initial system parameters. If you do not pass a config file, the implemented defaults are used. Note that you currently cannot change the geometry of the systems in your state once they are initialized.

State manipulation		Returns
setup(configfile="",	quiet=False)	None
delete(p_state)		None

7.2 Chain

For having more images one can copy the active image in the Clipboard and then insert in a specified position of the chain.

```
chain.Image_to_Clipboard(p_state )  # Copy p_state to Clipboard
chain.Insert_Image_After(p_state )  # Insert the image from Clipboard right after_

→ the currently active image
```

For getting the total number of images in the chain

```
number_of_images = chain.Get_NOI(p_state )
```

Get Info	Returns	Description
Get_Index(p_state)	int	Get Chain index
Get_NOI(p_state, idx_chain=-1)	int	Get Chain number of images
Get_Rx(p_state, idx_chain=-1)	"Array "	Get Rx
<pre>Get_Rx_Interpolated(p_state, idx_c hain=-1)</pre>	"Array (float) "	Get Rx interpolated
"Get_Energy(p_state, idx_chain=-1)"	"Array (float) "	Get Energy of every System in Chain
<pre>Get_Energy_Interpolated(p_state, i dx_chain=-1)</pre>	"Array (float) "	Get interpolated Energy of every System in Chain

Image Manipulation	Re- Description turns
<pre>Next_Image(p_state, idx_chain=- 1)</pre>	"None Switch active to next image of chain (one with largest index). If the current active is the last there is no effect.
<pre>Prev_Image(p_state, idx_chain=- 1)</pre>	"None Switch active to previous image of chain (one with smaller index). If the current active is the first one there is no effect
<pre>Jump_To_Image(p_state, idx_imag e=-1, idx_chain=-1)</pre>	"None Switch active to specific image of chain. If this image does not exist there is no effect.
<pre>Image_to_Clipboard(p_state, idx _image=-1, idx_chain=-1)</pre>	"None" Copy active image to clipboard
Replace_Image(p_state, idx_imag e=-1, idx_chain=-1)	"None Replace active image in chain. If the image does not exist there is no effect.
<pre>Insert_Image_Before(p_state, id x_image=-1, idx_chain=-1)</pre>	'None' Inserts clipboard image before the current active image. Active image index is increment by one.
<pre>Insert_Image_After(p_state, idx _image=-1, idx_chain=-1)</pre>	"None Insert clipboard image after the current active image. Active image has the same index.
Push_Back(p_state, idx_chain=-1)	"None Insert clipboard image at end of chain (after the image with the largest index).
<pre>Delete_Image(p_state, idx_image =-1, idx_chain=-1)</pre>	Delete active image. If index is specified delete the corresponding image. If the image does not exist there is no effect.
"Pop_Back(p_state, idx_chain=-1)"	"None" Delete image at end of chain.

Data	Re-	Description
	turns	
Update_Data(p_state, idx	None	Update the chain's data (interpolated energies
_chain=-1)		etc.)
Setup_Data(p_state, idx_	None	Setup the chain's data arrays
chain=-1)		

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7.3 System

System	Ret urn s	Description
Get_Index(p_state)	"i nt"	Returns the index of the currently active image
<pre>Get_NOS(p_state, idx_image=- 1, idx_chain=-1)</pre>	"i nt"	Returns the number of spins
"Get_Spin_Directions(p_state, idx_image=-1, idx_chain=-1)"	"[3*N OS] "	Returns an numpy. Array of size 3*NOS with the components of each spin's vector
<pre>Get_Energy(p_state, idx_imag e=-1, idx_chain=-1)</pre>	f loa t	Returns the energy of the system
<pre>Update_Data(p_state, idx_ima ge=-1, idx_chain=-1)</pre>	"N one "	Update the data of the state
Print_Energy_Array(p_statidx_image=-1,idx_chain=-1)	e ''N one ''	Print the energy array of the state

7.4 Constants

Physical Constants	Returns	Description
mu_B()	float	The Bohr Magneton [meV / T]
k_B()	float	The Boltzmann constant [meV / K]
hbar()	float	Planck's constant over 2pi [meV*ps / rad]
mRy()	float	Millirydberg [mRy / meV]
gamma()	float	The Gyromagnetic ratio of electron [rad / (ps*T)]
g_e()	float	The Electron g-factor [unitless]

7.5 Geometry

Get Geometry parameters	Returns	Description
Get_Bounds(p_state,	(([0], [0]))	Get bounds (minimum and maxi-
idx_image=-1 ,	<mark>"[3], [3]''</mark>	mum arrays)
idx_chain=-1)		
Get_Center(p_state,	66 Gard Gard Gard 6	Get center
idx_image=-1 ,	"float, fl oat, float"	
idx_chain=-1)		
Get_Basis_Vectors(p_state	,[3],[3],[3]	Get basis vectors
idx_i mage=-1,		
idx_chain=-1)		
<pre>Get_N_Cells(p_state,</pre>	"Int, Int, Int"	Get number of cells in each dimen-
idx_image=- 1,	Int, 111t, 111t	sion
idx_chain=-1)		
"Get_Translation_Vectors(p_state,	[3],[3],[3]	Get translation vectors
idx_image=-1, idx_chain=-		
1)"		
Get_Dimensionality(p_stat	eint	Get dimensionality of the system
idx_ image=-1,		
idx_chain=-1)		
Get_Spin_Positions(p_stat	e,[3*NOS]	Get Spin positions
idx_ image=-1,		
idx_chain=-1)		
<pre>Get_Atom_Types(p_state,</pre>	[NOS]	Get atom types
$idx_imag e=-1$,		
idx_chain=-1)		

7.6 Hamiltonian

Set Parameters	Re-	Description
	turns	
Set_Field(p_state, magnitude, direc tion,	None	Set external mag-
idx_image=-1, idx_chain=-1)		netic field
Set_Anisotropy(p_state, magnitude, direction,	None	Set anisotropy
idx_image=-1, idx_chain=-1)		

7.7 Log

Log manipulation	Return	Description
	s	
<pre>Send(p_state, level, sender, message, idx_ima ge=-1, idx_chain=-1)</pre>	"None	Send a Log message
Append(p_state)	"None	Append Log to file

7.5. Geometry 43

7.8 Parameters

7.8.1 LLG

Set LLG Parameters	Re-
	turns
setIterations(p_state, n_iterations, n_iterations_log, idx_i mage=-1,	None
idx_chain=-1)	
setDirectMinimization(p_state, use_minimization, idx_image=- 1,	None
idx_chain=-1)	
<pre>setConvergence(p_state, convergence, idx_image=-1, idx_chain =-1)</pre>	None
<pre>setTimeStep(p_state, dt, idx_image=-1, idx_chain=-1)</pre>	None
setDamping(p_state, damping, idx_image=-1, idx_chain=-1)	None
<pre>setSTT(p_state, use_gradient, magnitude, direction, idx_imag e=-1,</pre>	None
idx_chain=-1)	
<pre>setTemperature(p_state, temperature, idx_image=-1, idx_chain =-1)</pre>	None

Get LLG Parameters	Returns
<pre>getIterations(p_state, idx_image=-1, idx_chain=-1)</pre>	int, int
<pre>getDirect_Minimization(p_state, idx_image=-1,</pre>	int
idx_chain=-1)	
<pre>getConvergence(p_state, idx_image=-1, idx_chain=-1)</pre>	float
<pre>getTimeStep(p_state, idx_image=-1, idx_chain=-1)</pre>	float
<pre>getDamping(p_state, idx_image=-1, idx_chain=-1)</pre>	float
<pre>getSTT(p_state, idx_image=-1, idx_chain=-1)</pre>	float, [3],
	bool
<pre>getTemperature(p_state, idx_image=-1, idx_chain=-1)</pre>	float

7.8.2 **GNEB**

Set GNEB Parameters	Re-
	turns
setIterations(p_state, n_iterations, n_iterations_log, idx_im age=-1,	None
idx_chain=-1)	
setConvergence(p_state, convergence, idx_image=-1, idx_chain= -1)	None
setSpringConstant(p_state, c_spring, idx_image=-1, idx_chain= -1)	None
<pre>setClimbingFalling(p_state, image_type, idx_image=-1, idx_cha in=-1)</pre>	None
<pre>setImageTypeAutomatically(p_state, idx_chain=-1)</pre>	None

Get GNEB Parameters	Returns
<pre>getIterations(p_state, idx_chain=-1)</pre>	int, int
<pre>getConvergence(p_state, idx_image=-1, idx_chain=-1)</pre>	float
<pre>getSpringConstant(p_state, idx_image=-1, idx_chain=-1)</pre>	float
<pre>getClimbingFalling(p_state, idx_image=-1, idx_chain=-1)</pre>	int
<pre>getEnergyInterpolations(p_state, idx_chain=-1)</pre>	int

7.9 Quantities

Get Physical Quantities			Returns
<pre>Get_Magnetization(p_state,</pre>	idx_image=-1,	idx_chain=-1)	[3*float]

7.10 Simulation

The available method_types are:

Method	Argument
Landau-Lifshitz-Gilbert	"LLG"
Geodesic Nudged Elastic Band	"GNEB"
Monte-Carlo	"MC"

The available solver_types are:

Solver	Argument
Semi-Implicit Method B	"SIB"
Heun Method	"Heun"
Depondt Method	"Depondt"
Velocity Projection	"VP"
Nonlinear Conjugate Gradient	"NCG"

Note that the VP and NCG Solvers are only meant for direct minimization and not for dynamics.

Simulation state	Retur
	ns
SingleShot(p_state, method_type, solver_type, n_iterations=-1, n	Non e
_iterations_log=-1, idx_image=-1, idx_chain=-1)	
PlayPause(p_state, method_type, solver_type, n_iterations=-1, n_	Non e
iterations_log=-1, idx_image=-1, idx_chain=-1)	
Stop_All(p_state)	Non e
Running_Image(p_state, idx_image=-1, idx_chain=-1)	"Boo
	lean''
Dunning Chair (n. state idy shair 1)	ican
Running_Chain(p_state, idx_chain=-1)	" Boo
	lean''
Running_Collection(p_state)	66D = =
	"Boo
	lean''
Running_Anywhere_Chain(p_state, idx_chain=-1)	"Boo
	lean''
Running Anywhere Collection(p state)	
3_ 1	"Boo
	lean''

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7.11 Transition

Transition options	Ret urn s	Description
Homogeneous(p_state,	(0)	Generate homogeneous transition
idx_1, id x_2,	"N one "	between two images of a chain
idx_chain=-1)		
"Add_Noise_Temperature(p_state,	"N one "	Add some temperature-scaled noise to a transition between two images
temperature, idx_1, idx_2,		of a chain
idx_		of a chain
chain=-1)"		

7.12 Input/Output

Macros of File Formats for Vector Fields	values	Description
IO_Fileformat_Regular	0	sx sy sz (separated by whitespace)
IO_Fileformat_Regular_Pos	1	px py pz sx sy sz (separated by whitespace)
IO_Fileformat_CSV	2	sx, sy, sz (separated by commas)
IO_Fileformat_CSV_Pos	3	px, py, pz, sx, sy, (sz separated by commas)
IO_Fileformat_OVF_bin8	4	OOMMF vector field (OVF) v2.0 file format
IO_Fileformat_OVF_text	6	

For Image	Description
<pre>Image_Read(p_state, filename, fileformat=0, idx_i mage=-1,</pre>	Read an image from
idx_chain=-1)	disk
<pre>Image_Write(p_state, filename, fileformat=0, comm ent=" ",</pre>	Write an image to disk
idx_image=-1, idx_chain=-1)	
<pre>Image_Append(p_state, filename, fileformat=0, com ment="</pre>	Append an image to an
", idx_image=-1, idx_chain=-1)	existing file

For Chain	Description
Chain_Read(p_state, filename, fileformat=0, id	Read a chain of images
x_chain=-1)	from disk
Chain_Write(p_state, filename, fileformat=0, c omment="	Write a chain of images to
", idx_chain=-1)	disk
Chain_Append(p_state, filename, fileformat=0, comment="	Append a chain of images
", idx_chain=-1)	to disk

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CHAPTER 8

Contributors

8.1 Gideon Müller

- RWTH Aachen
- · University of Iceland
- PGI-1/IAS-1 at Forschungszentrum Jülich

General code design and project setup (including CMake). Implementation of the core library and UIs, most notably: - GNEB and MMF methods - Velocity projection solver - CUDA and OpenMP parallelizations of backend - C API and Pyhton bindings - C++ QT GUI and initial OpenGL code

(Oct. 2014 - ongoing)

8.2 Daniel Schürhoff

- RWTH Aachen
- PGI-1/IAS-1 at Forschungszentrum Jülich

Implementation of the initial core library, notably translating from Fortran90 to C++ and addition of STT to the SIB solver. Work on QT GUI and Python bindings.

(Oct. 2015 - Sept. 2016)

8.3 Nikolai Kiselev

• PGI-1/IAS-1 at Forschungszentrum Jülich

Scientific advice, general help and feedback, initial (Fortran90) implementations of: - isotropic Heisenberg Hamiltonian - Neighbour calculations - SIB solver - Monte Carlo methods

(2007 - ongoing)

8.4 Florian Rhiem

• Scientific IT-Systems, PGI/JCNS at Forschungszentrum Jülich

Implementation of C++ OpenGL code (VFRendering library), as well as JavaScript Web UI and WebGL code. Code design improvements, including the C API and CMake.

(Jan. 2016 - ongoing)

8.5 Stefanos Mayros

· RWTH Aachen

Work on unit testing and documentation and implementation of the Depondt solver.

(April 2017 - ongoing)

8.6 Constantin Disselkamp

• RWTH Aachen

Implementation and testing of gradient approximation of spin transfer torque.

(April 2017 - Juli 2017)

8.7 Pavel Bessarab

• ...

Help with the initial GNEB implementation. Initial implementation of the HTST method.

(April 2015 - ongoing)

8.8 Ingo Heimbach

• Scientific IT-Systems, PGI/JCNS at Forschungszentrum Jülich

Implementation of the initial OpenGL code. Code design suggestions and other general help.

(Jan. 2016 - ongoing)

8.9 Mathias Redies, Maximilian Merte, Rene Suckert

· RWTH Aachen

Initial CUDA implementation and tests. Code optimizations, suggestions and feedback.

(Sept. 2016 - Dec. 2016)

8.10 David Bauer

- RWTH Aachen
- PGI-1/IAS-1 at Forschungszentrum Jülich

Initial (Fortran90) implementations of the isotropic Heisenberg Hamiltonian, Neighbour calculations and the SIB solver.

(Oct. 2007 - Sept. 2008)

8.11 Graph

You may also take a look at the contributors graph.

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CHAPTER 9

Reference

The **Spirit** framework is a scientific project. If you present and/or publish scientific results or visualisations that used Spirit, you should add a reference.

9.1 The Framework

If you used any components of this framework please add a reference to our GitHub page. You may use e.g. the following TeX code:

```
\bibitem{spirit}
{Spirit spin simulation framework} (see spirit-code.github.io)
```

9.2 Specific Methods

The following need only be cited if used.

Depondt Solver

This Heun-like method for solving the LLG equation including the stochastic term has been published by Depondt et al.: http://iopscience.iop.org/0953-8984/21/33/336005 You may use e.g. the following TeX code:

```
\bibitem{Depondt}
Ph. Depondt et al. \textit{J. Phys. Condens. Matter} \textbf{21}, 336005 (2009).
```

SIB Solver

This stable method for solving the LLG equation efficiently and including the stochastic term has been published by Mentink et al.: http://iopscience.iop.org/0953-8984/22/17/176001 You may use e.g. the following TeX code:

```
\bibitem{SIB}
J. H. Mentink et al. \textit{J. Phys. Condens. Matter} \textbf{22}, 176001 (2010).
```

VP Solver

This intuitive direct minimization routine has been published as supplementary material by Bessarab et al.: http://www.sciencedirect.com/science/article/pii/S0010465515002696 You may use e.g. the following TeX code:

```
\bibitem{VP}
P. F. Bessarab et al. \textit{Comp. Phys. Comm.} \textbf{196}, 335 (2015).
```

GNEB Method

This specialized nudged elastic band method for calculating transition paths of spin systems has been published by Bessarab et al.: http://www.sciencedirect.com/science/article/pii/S0010465515002696 You may use e.g. the following TeX code:

```
\bibitem{GNEB}
P. F. Bessarab et al. \textit{Comp. Phys. Comm.} \textbf{196}, 335 (2015).
```

9.3 Papers

Here we will list select papers for which the Spirit framework was used, which may be of interest to you.

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Included Dependencies

10.1 Vector Field Rendering

libvfrendering is a C++ library for rendering vectorfields using OpenGL. Originally developed for spirit and based on WegGLSpins.js, it has an extendable architecture and currently offers renderer implementations for: - glyph-based vector field representations as arrows - colormapped surface and isosurface rendering - mapping vector directions onto a sphere

The library is still very much a work-in-progress, so its API is not yet stable and there are still several features missing that will be added in later releases. If you miss a feature or have another idea on how to improve libvfrendering, please open an issue or pull request!

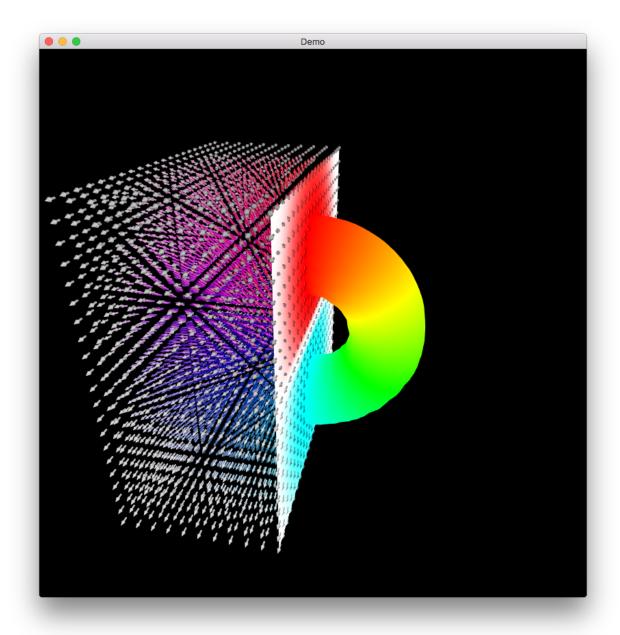
10.1.1 Getting Started

To use **libvfrendering**, you need to perform the following steps:

- 1. Include < VFRendering/View.hxx>
- 2. Create a VFRendering::Geometry
- 3. Read or calculate the vector directions
- 4. Pass geometry and directions to a VFRendering::View
- 5. Draw the view in an existing OpenGL context

1. Include <VFRendering/View.hxx>

When using **libvfrendering**, you will mostly interact with View objects, so it should be enough to #include <VFRendering/View.hxx>.



2. Create a VFRendering::Geometry

The **geometry describes the positions** on which you evaluated the vector field and how they might form a grid (optional, e.g. for isosurface and surface rendering). You can pass the positions directly to the constructor or call one of the class' static methods.

As an example, this is how you could create a simple, cartesian 30x30x30 geometry, with coordinates between -1 and 1.

3. Read or calculate the vector directions

This step highly depends on your use case. The **directions are stored as a "std::vector<glm::vec3>"**, so they can be created in a simple loop:

As shown here, the directions should be in \mathbf{C} order when using the VFRendering::Geometry static methods. If you do not know glm, think of a glm::vec3 as a struct containing three floats x, y and z.

4. Create a VFRendering::VectorField

This class simply contains geometry and directions.

```
VFRendering::VectorField vf(geometry, directions);
```

To update the VectorField data, use VectorField::update. If the directions changed but the geometry is the same, you can use the VectorField::updateVectors method or VectorField::updateGeometry vice versa.

5. Create a VFRendering::View and a Renderer

The view object is what you will interact most with. It provides an interface to the various renderers and includes functions for handling mouse input.

You can **create a new view** and then **initialize the renderer(s)** (as an example, we use the VFRendering::ArrowRenderer):

```
VFRendering::View view;
auto arrow_renderer_ptr = std::make_shared<VFRendering::ArrowRenderer>(view, vf);
view.renderers( {{ arrow_renderer_ptr, {0, 0, 1, 1} }});
```

5. Draw the view in an existing OpenGL context

To actually see something, you need to create an OpenGL context using a toolkit of your choice, e.g. Qt or GLFW. After creating the context, pass the framebuffer size to the **setFramebufferSize method**. You can then call the **draw method** of the view to render the vector field, either in a loop or only when you update the data.

```
view.draw();
```

For a complete example, including an interactive camera, see demo.cxx.

10.1.2 Python Package

The Python package has bindings which correspond directly to the C++ class and function names. To use **pyVFRendering**, you need to perform the following steps:

- 1. import pyVFRendering as vfr
- 2. Create a vfr. Geometry
- 3. Read or calculate the vector directions
- 4. Pass geometry and directions to a vfr. View
- 5. Draw the view in an existing OpenGL context

1. import

In order to import pyVFRendering as vfr, you can either pip install pyVFRendering or download and build it yourself.

You can build with python3 setup.py build, which will generate a library somewhere in your build subfolder, which you can import in python. Note that you may need to add the folder to your PYTHONPATH.

2. Create a pyVFRendering.Geometry

As above:

3. Read or calculate the vector directions

This step highly depends on your use case. Example:

```
directions = []
for iz in range(n_cells[2]):
    for iy in range(n_cells[1]):
        for ix in range(n_cells[0]):
            # calculate direction for ix, iy, iz
            directions.append([ix-4.5, iy-4.5, iz-4.5])
```

4. Pass geometry and directions to a pyVFRendering. View

You can create a new view and then pass the geometry and directions by calling the update method:

```
view = vfr.View()
view.update(geometry, directions)
```

If the directions changed but the geometry is the same, you can use the **updateVectors method**.

5. Draw the view in an existing OpenGL context

To actually see something, you need to create an OpenGL context using a framework of your choice, e.g. Qt or GLFW. After creating the context, pass the framebuffer size to the **setFramebufferSize method**. You can then call the **draw method** of the view to render the vector field, either in a loop or only when you update the data.

For a complete example, including an interactive camera, see demo.py.

10.1.3 Renderers

libvfrendering offers several types of renderers, which all inherit from VFRendering::RendererBase. The most relevant are the VectorFieldRenderers:

- VFRendering::ArrowRenderer, which renders the vectors as colored arrows
- VFRendering::SphereRenderer, which renders the vectors as colored spheres
- VFRendering::SurfaceRenderer, which renders the surface of the geometry using a colormap
- VFRendering::IsosurfaceRenderer, which renders an isosurface of the vectorfield using a colormap
- VFRendering::VectorSphereRenderer, which renders the vectors as dots on a sphere, with the position of each dot representing the direction of the vector

In addition to these, there also the following renderers which do not require a VectorField: - VFRendering::CombinedRenderer, which can be used to create a combination of several renderers, like an isosurface rendering with arrows - VFRendering::BoundingBoxRenderer, which is used for rendering bounding boxes around the geometry rendered by an VFRendering::ArrorRenderer, VFRendering::SurfaceRenderer or VFRendering::IsosurfaceRenderer - VFRendering::CoordinateSystemRenderer, which is used for rendering a coordinate system, with the axes colored by using the colormap

To control what renderers are used, you can use VFRendering::View::renderers, where you can pass it a std::vectors of std::pairs of renderers as std::shared_ptr<VFRendering::RendererBase> (i.e. shared pointers) and viewports as glm::vec4.

10.1.4 Options

To modify the way the vector field is rendered, **libvfrendering** offers a variety of options. To set these, you can create an **VFRendering::Options** object.

As an example, to adjust the vertical field of view, you would do the following:

```
VFRendering::Options options;
options.set<VFRendering::View::Option::VERTICAL_FIELD_OF_VIEW>(30);
view.updateOptions(options);
```

If you want to set only one option, you can also use **View::setOption**:

```
view.setOption<VFRendering::View::Option::VERTICAL_FIELD_OF_VIEW>(30);
```

If you want to set an option for an individual Renderer, you can use the methods **RendererBase::updateOptions** and **RendererBase::setOption** in the same way.

Whether this way of setting options should be replaced by getters/setters will be evaluated as the API becomes more stable.

Currently, the following options are available:

Index	Туре	Default value
View::Op tion::BO UNDING_ BOX_MIN	glm::vec 3	{-1, -1, -1}
View::Op tion::BO UNDING_ BOX_MAX	glm::vec 3	{1, 1, 1}
View::Op tion::SY STEM_CE NTER	glm::vec 3	$\{0, 0, 0\}$
View::Op tion::VE RTICAL_FIELD_O F_VIEW	float	45.0
View::Op tion::BA CKGROUND _COLOR	glm::vec 3	$\{0, 0, 0\}$
View::Op tion::CO LORMAP_ IMPLEMEN TATION	std::str ing	VFRendering::Uti lities::getColor map
View::Op tion::IS _VISIBL E_IMPLE MENTATIO N	std::str ing	bool is_visible(vec3 position, vec3 dir
View::Op tion::CA MERA_PO SITION	glm::vec 3	{14.5, 14.5, 30}
View::Op tion::CE NTER_PO SITION	glm::vec 3	{14.5, 14.5, 0}
View::Op tion::UP _VECTOR	glm::vec 3	{0, 1, 0}
ArrowRen derer::O ption::C ONE_RAD IUS	float	0.25
ArrowRen derer::O ption::C ONE_HEI GHT	float	0.6
ArrowRen derer::O ption::C YLINDER_RADIUS	float	0.125
ArrowRen derer::O ption::C YLINDER_HEIGHT	float	0.7
ArrowRen derer::O ption::L EVEL_OF _DETAIL	unsigned int	20
Bounding BoxRende rer::Opt ion::COL OR	glm::vec 3	{1.0, 1.0, 1.0}
Coordina teSystem Renderer ::Option ::AXIS_ LENGTH	glm::vec 3	{0.5, 0.5, 0.5}
Coordina teSystem Renderer ::Option ::ORIGIN	glm::vec 3	$\{0.0, 0.0, 0.0\}$
Coordina teSystem Renderer ::Option ::NORMAL IZE	bool	false
Coordina teSystem Renderer ::Option ::LEVEL_OF_DET AIL	unsigned int	100
Coordina teSystem Renderer ::Option ::CONE_ HEIGHT	float	0.3
Coordina teSystem Renderer ::Option ::CONE_ RADIUS	float	0.1
Coordina teSystem Renderer ::Option ::CYLIND ER_HEIG HT	float	0.7
Coordina teSystem Renderer ::Option ::CYLIND ER_RADI US	float	0.07
Isosurfa ceRender er::Opti on::ISOV ALUE	float	0.0
Isosurfa ceRender er::Opti on::LIGH TING_IM PLEMENTA TION	std::str ing	float lighting(vec3 position, vec3 norm
Isosurfa ceRender er::Opti on::VALU E_FUNCT ION	std::fun ction	[] (const glm::vec3& position, const g
VectorSp hereRend erer::Op tion::PO INT_SIZ E_RANGE	glm::vec 2	{1.0, 4.0}
VectorSp hereRend erer::Op tion::IN NER_SPH ERE_RAD IUS	float	0.95
VectorSp hereRend erer::Op tion::US E_SPHER E_FAKE_PERSPEC TIVE	bool	true

10.1.5 ToDo

- A EGS plugin for combining libvfrendering with existing EGS plugins.
- Methods for reading geometry and directions from data files

• Documentation

See the issues for further information and adding your own requests.

10.2 Eigen

Eigen is a C++ template library for linear algebra: matrices, vectors, numerical solvers, and related algorithms.

For more information go to http://eigen.tuxfamily.org/.

10.3 Spectra

Spectra stands for **Sp**arse Eigenvalue Computation Toolkit as a **R**edesigned ARPACK. It is a C++ library for large scale eigenvalue problems, built on top of Eigen, an open source linear algebra library.

Spectra is implemented as a header-only C++ library, whose only dependency, **Eigen**, is also header-only. Hence **Spectra** can be easily embedded in C++ projects that require calculating eigenvalues of large matrices.

10.3.1 Relation to ARPACK

ARPACK is a software written in FORTRAN for solving large scale eigenvalue problems. The development of **Spectra** is much inspired by ARPACK, and as the whole name indicates, **Spectra** is a redesign of the ARPACK library using C++ language.

In fact, **Spectra** is based on the algorithms described in the ARPACK Users' Guide, but it does not use the ARPACK code, and it is **NOT** a clone of ARPACK for C++. In short, **Spectra** implements the major algorithms in ARPACK, but **Spectra** provides a completely different interface, and it does not depend on ARPACK.

10.3.2 Common Usage

Spectra is designed to calculate a specified number (k) of eigenvalues of a large square matrix (A). Usually k is much less than the size of matrix (n), so that only a few eigenvalues and eigenvectors are computed, which in general is more efficient than calculating the whole spectral decomposition. Users can choose eigenvalue selection rules to pick up the eigenvalues of interest, such as the largest k eigenvalues, or eigenvalues with largest real parts, etc.

To use the eigen solvers in this library, the user does not need to directly provide the whole matrix, but instead, the algorithm only requires certain operations defined on A, and in the basic setting, it is simply the matrix-vector multiplication. Therefore, if the matrix-vector product $A \times C$ can be computed efficiently, which is the case when A is sparse, **Spectra** will be very powerful for large scale eigenvalue problems.

There are two major steps to use the **Spectra** library:

- 1. Define a class that implements a certain matrix operation, for example the matrix-vector multiplication y = A * x or the shift-solve operation $y = inv(A \sigma * I) * x$. **Spectra** has defined a number of helper classes to quickly create such operations from a matrix object. See the documentation of DenseGenMatProd, DenseSymShiftSolve, etc.
- 2. Create an object of one of the eigen solver classes, for example SymEigsSolver for symmetric matrices, and GenEigsSolver for general matrices. Member functions of this object can then be called to conduct the computation and retrieve the eigenvalues and/or eigenvectors.

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Below is a list of the available eigen solvers in **Spectra**: - SymEigsSolver: For real symmetric matrices - GenEigsSolver: For general real matrices - SymEigsShiftSolver: For real symmetric matrices using the shift-and-invert mode - GenEigsRealShiftSolver: For general real matrices using the shift-and-invert mode, with a real-valued shift - GenEigsComplexShiftSolver: For general real matrices using the shift-and-invert mode, with a complex-valued shift - SymGEigsSolver: For generalized eigen solver for real symmetric matrices

10.3.3 Examples

Below is an example that demonstrates the use of the eigen solver for symmetric matrices.

```
#include <Eigen/Core>
#include <SymEigsSolver.h> // Also includes <MatOp/DenseSymMatProd.h>
#include <iostream>
using namespace Spectra;
int main()
    // We are going to calculate the eigenvalues of M
   Eigen::MatrixXd A = Eigen::MatrixXd::Random(10, 10);
   Eigen::MatrixXd M = A + A.transpose();
    // Construct matrix operation object using the wrapper class DenseGenMatProd
    DenseSymMatProd<double> op (M);
    // Construct eigen solver object, requesting the largest three eigenvalues
    SymEigsSolver< double, LARGEST_ALGE, DenseSymMatProd<double> > eigs(&op, 3, 6);
   // Initialize and compute
   eigs.init();
   int nconv = eigs.compute();
   // Retrieve results
   Eigen::VectorXd evalues;
    if(eigs.info() == SUCCESSFUL)
        evalues = eigs.eigenvalues();
    std::cout << "Eigenvalues found:\n" << evalues << std::endl;</pre>
    return 0;
```

Sparse matrix is supported via the SparseGenMatProd class.

```
#include <Eigen/Core>
#include <Eigen/SparseCore>
#include <GenEigsSolver.h>
#include <MatOp/SparseGenMatProd.h>
#include <iostream>
using namespace Spectra;
int main()
{
    // A band matrix with 1 on the main diagonal, 2 on the below-main subdiagonal,
    // and 3 on the above-main subdiagonal
    const int n = 10;
```

```
Eigen::SparseMatrix<double> M(n, n);
M.reserve(Eigen::VectorXi::Constant(n, 3));
for (int i = 0; i < n; i++)
    M.insert(i, i) = 1.0;
    if(i > 0)
        M.insert(i - 1, i) = 3.0;
    if(i < n - 1)
        M.insert(i + 1, i) = 2.0;
}
// Construct matrix operation object using the wrapper class SparseGenMatProd
SparseGenMatProd<double> op (M);
// Construct eigen solver object, requesting the largest three eigenvalues
GenEigsSolver< double, LARGEST_MAGN, SparseGenMatProd<double> > eigs(&op, 3, 6);
// Initialize and compute
eigs.init();
int nconv = eigs.compute();
// Retrieve results
Eigen::VectorXcd evalues;
if(eigs.info() == SUCCESSFUL)
    evalues = eigs.eigenvalues();
std::cout << "Eigenvalues found:\n" << evalues << std::endl;</pre>
return 0;
```

And here is an example for user-supplied matrix operation class.

```
#include <Eigen/Core>
#include <SymEigsSolver.h>
#include <iostream>
using namespace Spectra;
// M = diag(1, 2, ..., 10)
class MyDiagonalTen
public:
   int rows() { return 10; }
   int cols() { return 10; }
   // y_out = M * x_in
   void perform_op(double *x_in, double *y_out)
        for(int i = 0; i < rows(); i++)</pre>
            y_{out[i]} = x_{in[i]} * (i + 1);
    }
};
int main()
   MyDiagonalTen op;
```

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```
SymEigsSolver<double, LARGEST_ALGE, MyDiagonalTen> eigs(&op, 3, 6);
eigs.init();
eigs.compute();
if(eigs.info() == SUCCESSFUL)
{
    Eigen::VectorXd evalues = eigs.eigenvalues();
    std::cout << "Eigenvalues found:\n" << evalues << std::endl;
}
return 0;
}</pre>
```

10.3.4 Shift-and-invert Mode

When we want to find eigenvalues that are closest to a number σ , for example to find the smallest eigenvalues of a positive definite matrix (in which case $\sigma = 0$), it is advised to use the shift-and-invert mode of eigen solvers.

In the shift-and-invert mode, selection rules are applied to $1/(\lambda - \sigma)$ rather than λ , where λ are eigenvalues of A. To use this mode, users need to define the shift-solve matrix operation. See the documentation of SymEigsShiftSolver for details.

10.3.5 Documentation

The API reference page contains the documentation of **Spectra** generated by Doxygen, including all the background knowledge, example code and class APIs.

More information can be found in the project page http://yixuan.cos.name/spectra.

10.3.6 License

Spectra is an open source project licensed under MPL2, the same license used by **Eigen**.

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Documentation

10.4.1 Features

- Two APIs: faster concatenation-based write API and slower, but still very fast, replacement-based format API with positional arguments for localization.
- Write API similar to the one used by IOStreams but stateless allowing faster implementation.
- Format API with format string syntax similar to the one used by str.format in Python.
- Safe printf implementation including the POSIX extension for positional arguments.
- Support for user-defined types.

- High speed: performance of the format API is close to that of glibc's printf and better than the performance of IOStreams. See *Speed tests* and Fast integer to string conversion in C++.
- Small code size both in terms of source code (the core library consists of a single header file and a single source file) and compiled code. See *Compile time and code bloat*.
- Reliability: the library has an extensive set of unit tests.
- Safety: the library is fully type safe, errors in format strings are reported using exceptions, automatic memory management prevents buffer overflow errors.
- Ease of use: small self-contained code base, no external dependencies, permissive BSD license
- Portability with consistent output across platforms and support for older compilers.
- Clean warning-free codebase even on high warning levels (-Wall -Wextra -pedantic).
- Support for wide strings.
- Optional header-only configuration enabled with the FMT_HEADER_ONLY macro.

See the documentation for more details.

10.4.2 Examples

This prints Hello, world! to stdout:

```
fmt::print("Hello, {}!", "world"); // uses Python-like format string syntax
fmt::printf("Hello, %s!", "world"); // uses printf format string syntax
```

Arguments can be accessed by position and arguments' indices can be repeated:

```
std::string s = fmt::format("{0}{1}{0}", "abra", "cad");
// s == "abracadabra"
```

fmt can be used as a safe portable replacement for itoa:

An object of any user-defined type for which there is an overloaded std::ostream insertion operator (operator<<) can be formatted:

```
#include "fmt/ostream.h"

class Date {
  int year_, month_, day_;
  public:
   Date(int year, int month, int day) : year_(year), month_(month), day_(day) {}

  friend std::ostream &operator<<(std::ostream &os, const Date &d) {
    return os << d.year_ << '-' << d.month_ << '-' << d.day_;
  }
};

std::string s = fmt::format("The date is {}", Date(2012, 12, 9));
// s == "The date is 2012-12-9"</pre>
```

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You can use the FMT_VARIADIC macro to create your own functions similar to format and print which take arbitrary arguments:

```
// Prints formatted error message.
void report_error(const char *format, fmt::ArgList args) {
  fmt::print("Error: ");
  fmt::print(format, args);
}
FMT_VARIADIC(void, report_error, const char *)
report_error("file not found: {}", path);
```

Note that you only need to define one function that takes fmt::ArgList argument. FMT_VARIADIC automatically defines necessary wrappers that accept variable number of arguments.

10.4.3 Projects using this library

- 0 A.D.: A free, open-source, cross-platform real-time strategy game
- AMPL/MP: An open-source library for mathematical programming
- CUAUV: Cornell University's autonomous underwater vehicle
- Drake: A planning, control, and analysis toolbox for nonlinear dynamical systems (MIT)
- Envoy: C++ L7 proxy and communication bus (Lyft)
- FiveM: a modification framework for GTA V
- HarpyWar/pvpgn: Player vs Player Gaming Network with tweaks
- KBEngine: An open-source MMOG server engine
- Keypirinha: A semantic launcher for Windows
- Kodi (formerly xbmc): Home theater software
- Lifeline: A 2D game
- MongoDB Smasher: A small tool to generate randomized datasets
- OpenSpace: An open-source astrovisualization framework
- PenUltima Online (POL): An MMO server, compatible with most Ultima Online clients
- quasardb: A distributed, high-performance, associative database
- readpe: Read Portable Executable
- · redis-cerberus: A Redis cluster proxy
- Saddy: Small crossplatform 2D graphic engine
- Salesforce Analytics Cloud: Business intelligence software
- Scylla: A Cassandra-compatible NoSQL data store that can handle 1 million transactions per second on a single server
- Seastar: An advanced, open-source C++ framework for high-performance server applications on modern hardware
- spdlog: Super fast C++ logging library
- Stellar: Financial platform

- Touch Surgery: Surgery simulator
- TrinityCore: Open-source MMORPG framework

More...

If you are aware of other projects using this library, please let me know by email or by submitting an issue.

10.4.4 Motivation

So why yet another formatting library?

There are plenty of methods for doing this task, from standard ones like the printf family of function and IOStreams to Boost Format library and FastFormat. The reason for creating a new library is that every existing solution that I found either had serious issues or didn't provide all the features I needed.

Printf

The good thing about printf is that it is pretty fast and readily available being a part of the C standard library. The main drawback is that it doesn't support user-defined types. Printf also has safety issues although they are mostly solved with __attribute__ ((format (printf, ...)) in GCC. There is a POSIX extension that adds positional arguments required for i18n to printf but it is not a part of C99 and may not be available on some platforms.

IOStreams

The main issue with IOStreams is best illustrated with an example:

```
std::cout << std::setprecision(2) << std::fixed << 1.23456 << "\n";
```

which is a lot of typing compared to printf:

```
printf("%.2f\n", 1.23456);
```

Matthew Wilson, the author of FastFormat, referred to this situation with IOStreams as "chevron hell". IOStreams doesn't support positional arguments by design.

The good part is that IOStreams supports user-defined types and is safe although error reporting is awkward.

Boost Format library

This is a very powerful library which supports both printf-like format strings and positional arguments. The main its drawback is performance. According to various benchmarks it is much slower than other methods considered here. Boost Format also has excessive build times and severe code bloat issues (see *Benchmarks*).

FastFormat

This is an interesting library which is fast, safe and has positional arguments. However it has significant limitations, citing its author:

Three features that have no hope of being accommodated within the current design are:

- Leading zeros (or any other non-space padding)
- Octal/hexadecimal encoding

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• Runtime width/alignment specification

It is also quite big and has a heavy dependency, STLSoft, which might be too restrictive for using it in some projects.

Loki SafeFormat

SafeFormat is a formatting library which uses printf-like format strings and is type safe. It doesn't support user-defined types or positional arguments. It makes unconventional use of operator () for passing format arguments.

Tinyformat

This library supports printf-like format strings and is very small and fast. Unfortunately it doesn't support positional arguments and wrapping it in C++98 is somewhat difficult. Also its performance and code compactness are limited by IOStreams.

Boost Spirit.Karma

This is not really a formatting library but I decided to include it here for completeness. As IOStreams it suffers from the problem of mixing verbatim text with arguments. The library is pretty fast, but slower on integer formatting than fmt::Writer on Karma's own benchmark, see Fast integer to string conversion in C++.

10.4.5 Benchmarks

Speed tests

The following speed tests results were generated by building tinyformat_test.cpp on Ubuntu GNU/Linux 14.04.1 with g++-4.8.2 -03 -DSPEED_TEST -DHAVE_FORMAT, and taking the best of three runs. In the test, the format string "%0.10f:%04d:%+g:%s:%p:%c:%%\n" or equivalent is filled 20000000 times with output sent to /dev/null; for further details see the source.

Library	Method	Run Time, s
EGLIBC 2.19	printf	1.30
libstdc++ 4.8.2	std::ostream	1.85
fmt 1.0	fmt::print	1.42
tinyformat 2.0.1	tfm::printf	2.25
Boost Format 1.54	boost::format	9.94

As you can see boost::format is much slower than the alternative methods; this is confirmed by other tests. Tinyformat is quite good coming close to IOStreams. Unfortunately tinyformat cannot be faster than the IOStreams because it uses them internally. Performance of fmt is close to that of printf, being faster than printf on integer formatting, but slower on floating-point formatting which dominates this benchmark.

Compile time and code bloat

The script bloat-test.py from format-benchmark tests compile time and code bloat for nontrivial projects. It generates 100 translation units and uses printf() or its alternative five times in each to simulate a medium sized project. The resulting executable size and compile time (g++-4.8.1, Ubuntu GNU/Linux 13.10, best of three) is shown in the following tables.

Optimized build (-O3)

Method	Compile Time, s	Executable size, KiB	Stripped size, KiB
printf	2.6	41	30
IOStreams	19.4	92	70
fmt	46.8	46	34
tinyformat	64.6	418	386
Boost Format	222.8	990	923

As you can see, fmt has two times less overhead in terms of resulting code size compared to IOStreams and comes pretty close to printf. Boost Format has by far the largest overheads.

Non-optimized build

Method	Compile Time, s	Executable size, KiB	Stripped size, KiB
printf	2.1	41	30
IOStreams	19.7	86	62
fmt	47.9	108	86
tinyformat	27.7	234	190
Boost Format	122.6	884	763

libc, libstdc++ and libfmt are all linked as shared libraries to compare formatting function overhead only. Boost Format and tinyformat are header-only libraries so they don't provide any linkage options.

Running the tests

Please refer to Building the library for the instructions on how to build the library and run the unit tests.

Benchmarks reside in a separate repository, format-benchmarks, so to run the benchmarks you first need to clone this repository and generate Makefiles with CMake:

```
$ git clone --recursive https://github.com/fmtlib/format-benchmark.git
$ cd format-benchmark
$ cmake .
```

Then you can run the speed test:

```
$ make speed-test
```

or the bloat test:

```
$ make bloat-test
```

10.4.6 License

fmt is distributed under the BSD license.

The Format String Syntax section in the documentation is based on the one from Python string module documentation adapted for the current library. For this reason the documentation is distributed under the Python Software Foundation license available in doc/python-license.txt. It only applies if you distribute the documentation of fmt.

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10.4.7 Acknowledgments

The fmt library is maintained by Victor Zverovich (vitaut) and Jonathan Müller (foonathan) with contributions from many other people. See Contributors and Releases for some of the names. Let us know if your contribution is not listed or mentioned incorrectly and we'll make it right.

The benchmark section of this readme file and the performance tests are taken from the excellent tinyformat library written by Chris Foster. Boost Format library is acknowledged transitively since it had some influence on tinyformat. Some ideas used in the implementation are borrowed from Loki SafeFormat and Diagnostic API in Clang. Format string syntax and the documentation are based on Python's str.format. Thanks Doug Turnbull for his valuable comments and contribution to the design of the type-safe API and Gregory Czajkowski for implementing binary formatting. Thanks Ruslan Baratov for comprehensive comparison of integer formatting algorithms and useful comments regarding performance, Boris Kaul for C++ counting digits benchmark. Thanks to CarterLi for contributing various improvements to the code.

CHAPTER 11

Indices and tables

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